

# A Taste of the Standard Model

Lagrangian formalism

simple relation with Feynman rules

## Two Big Ideas :

★ Local gauge invariance

► Invariance of  $\mathcal{L}_{\text{fermion}}$  under transformations  
"requires" existence of massless vector bosons with specified interactions.

► Gauge invariant theories are renormalizable

★ Spontaneous symmetry breaking of scalar fields

The Higgs mechanism combines gauge invariance with spontaneous symmetry breaking to describe gauge invariant massive vector particles.

# Classical Lagrangians (review)

$$m \frac{d\vec{v}}{dt} = -\vec{\nabla} U \quad \text{equation of motion}$$

Lagrangian formulation

$$L = T - U \quad T = \frac{1}{2} m v^2$$

Euler-Lagrange equations:

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) = \frac{\partial L}{\partial q_i} \quad q = x, y, z$$

$$\text{so that } \frac{\partial L}{\partial \dot{q}_i} = m v_i \quad \frac{\partial L}{\partial q_i} = - \frac{\partial U}{\partial q_i}$$

$$\Rightarrow m \frac{dv_i}{dt} = - \frac{dU}{dq_i} \quad \text{as above}$$

# Field Theory Lagrangians

Instead of coordinates of an object we use fields that are functions of spacetime  $\phi_i(x^\mu)$

## Examples

- ▶  $\phi$  could represent temperature as a function of position (scalar field).
- ▶  $\phi_{i=1,2,3}$  could represent 3 components of an electric field  $\vec{E}$  (vector field).

generalized  
equ. of motion

$$\partial_\mu \left( \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \right) = \frac{\partial \mathcal{L}}{\partial \phi_i}$$

## Note

- ▶ space & time on equal footing
- ▶ prefer to work with Lagrangian

density  $\mathcal{L}$  such that  $L = \int d^3x \mathcal{L}$

In our case, the fields will be particle fields (i.e. wave functions).

# Lagrangian Densities for Particle Fields

Scalar (spin zero) field:

$$\mathcal{L} = \frac{1}{2}(\partial_\mu \phi)(\partial^\mu \phi) - \frac{1}{2} m^2 \phi^2$$

Check this gives the correct equ. of motion...

$$\frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} = \partial^\mu \phi \quad ; \quad \frac{\partial \mathcal{L}}{\partial \phi} = -m^2 \phi$$

Euler-Lagrange:  $\partial_\mu \partial^\mu \phi + m^2 \phi = 0$  ★

★ This is the Klein-Gordon equation for a spin zero relativistic particle.

Spin  $\frac{1}{2}$  field:

$$\mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi$$

Check:

$$\frac{\partial \mathcal{L}}{\partial(\partial_\mu \bar{\psi})} = 0 \quad \frac{\partial \mathcal{L}}{\partial \bar{\psi}} = i \gamma^\mu \partial_\mu \psi - m \psi$$

$$i \gamma^\mu \partial_\mu \psi - m \psi = 0 \quad \text{Dirac Equ.}$$

(satisfied by particle spinors)

$$\frac{\partial \mathcal{L}}{\partial(\partial_\mu \psi)} = i \bar{\psi} \gamma^\mu \quad \frac{\partial \mathcal{L}}{\partial \psi} = -m \bar{\psi}$$

$$i \partial_\mu \bar{\psi} \gamma^\mu + m \bar{\psi} = 0$$

(satisfied by antiparticle spinors)

photons (massless spin 1 fields):

First a reminder concerning the equations of motion. (section 7.4).

Maxwell :

$$\textcircled{i} \quad \nabla \cdot \bar{E} = 4\pi \rho \quad \textcircled{iii} \quad \nabla \cdot \bar{B} = 0$$

$$\textcircled{ii} \quad \nabla \times \bar{E} + \frac{1}{c} \frac{\partial \bar{B}}{\partial t} = 0 \quad \textcircled{iv} \quad \nabla \times \bar{B} - \frac{1}{c} \frac{\partial \bar{E}}{\partial t} = \frac{4\pi}{c} \bar{J}$$

define the field strength tensor:

$$F^{\mu\nu} = \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & -B_z & B_y \\ E_y & B_z & 0 & -B_x \\ E_z & -B_y & B_x & 0 \end{pmatrix}$$

i and iv imply ...

$$\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} J^\nu$$

where  $J^\nu = (c\rho, \vec{J})$

iii implies ...

1)  $\vec{B} = \nabla \times \vec{A}$  where  $\vec{A}$  is called the vector potential

and, using ii

$$\nabla \times \left( \vec{E} + \frac{1}{c} \frac{\partial \vec{A}}{\partial t} \right) = 0$$

2)  $\Rightarrow \vec{E} = -\nabla V - \frac{1}{c} \frac{\partial \vec{A}}{\partial t}$

where  $V$  is the scalar potential

1) and 2) lead to

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

where  $A^\mu = (V, \vec{A})$

Finally:  $\partial_\mu \partial^\mu A^\nu - \partial^\nu \partial_\mu A^\mu = \frac{4\pi}{c} J^\nu$

photon Lagrangian:

$$\mathcal{L} = \frac{-1}{16\pi} (\partial^\mu A^\nu - \partial^\nu A^\mu)(\partial_\mu A_\nu - \partial_\nu A_\mu) - J^\nu A_\nu$$

← convention

check:

$$\frac{\partial \mathcal{L}}{\partial(\partial_\mu A_\nu)} = \frac{-1}{4\pi} (\partial^\mu A^\nu - \partial^\nu A^\mu); \quad \frac{\partial \mathcal{L}}{\partial A_\nu} = -J^\nu$$

$$\Rightarrow \partial_\mu (\partial^\mu A^\nu - \partial^\nu A^\mu) = 4\pi J^\nu$$

also written  $\partial_\mu F^{\mu\nu} = 4\pi J^\nu$   
(Maxwell's equs., as we expected)

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A massive spin 1 free field gives

$$\mathcal{L} = -\frac{1}{16\pi} F^{\mu\nu} F_{\mu\nu} + \frac{1}{8\pi} m^2 A^\nu A_\nu$$

$$\partial_\mu F^{\mu\nu} + m^2 A^\nu = 0 \quad \text{Proca equation}$$

Incidentally,  $A^\mu$  is not uniquely determined.

$$\text{Let } A^{\mu'} = A^\mu + \partial^\mu \lambda$$

Then

$$\begin{aligned} \partial^\mu A^{\nu'} - \partial^\nu A^{\mu'} &= \partial^\mu (A^\nu + \partial^\nu \lambda) - \partial^\nu (A^\mu + \partial^\mu \lambda) \\ &= \partial^\mu A^\nu - \partial^\nu A^\mu \end{aligned}$$

$\Rightarrow F^{\mu\nu}$  and the equs. of motion are unchanged.

This is called a gauge transformation

# Gauge Invariance.

The Dirac Lagrangian:

$$\mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi$$

is invariant under the transformation

$$\psi \rightarrow e^{i\theta} \psi \quad (\text{and } \bar{\psi} \rightarrow e^{-i\theta} \bar{\psi})$$

This is called a "global gauge transformation"

Question:

What if the transformation were a function of spacetime?

$$\psi \rightarrow e^{i\theta(x)} \psi \quad x \text{ stands for } t, x, y, z$$

Then

$$\partial_\mu (e^{i\theta(x)} \psi) = i \underbrace{(\partial_\mu \theta(x))}_{\text{extra term}} e^{i\theta(x)} \psi + e^{i\theta(x)} \partial_\mu \psi$$

$$\mathcal{L} \rightarrow \mathcal{L} - (\partial_\mu \theta(x)) \bar{\psi} \gamma^\mu \psi$$

$\mathcal{L}$  is no longer invariant.

We want to try to restore the invariance by adding something to  $\mathcal{L}$ .

first, a notation change:

$$\lambda(x) \equiv \frac{-1}{g} \Theta(x)$$

$$\text{so } \mathcal{L} \rightarrow \mathcal{L} + g (\partial_\mu \lambda) \bar{\psi} \gamma^\mu \psi$$

now, the fix:

$$\text{try } \mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi - g \bar{\psi} \gamma^\mu \psi A_\mu$$

↑  
new gauge field

where the transformation

$$\text{rule for } A_\mu \text{ is } A_\mu \rightarrow A_\mu + \partial_\mu \lambda$$

as it was in Maxwell's equations.


The full Lagrangian is now invariant.


$\psi \rightarrow e^{-ig\lambda} \psi$  together with  $A_\mu \rightarrow A_\mu + \partial_\mu \lambda$  works!

This is called a "local gauge transformation"

BUT now we have introduced a new gauge field that couples to the fermions. We should also include some terms for the gauge particle itself.

$$\mathcal{L}_{\text{gauge}} = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} + \frac{1}{8\pi} m_A^2 A^\nu A^\mu$$


 this part is invariant  
 (as we saw for Maxwell's eqs.)


 this part is not  
 invariant

**conclude:** only massless vector gauge particles have local gauge invariant Lagrangians. ?

$$\mathcal{L}_{\text{total}} = \underbrace{[i\bar{\psi}\gamma^\mu\partial_\mu\psi - m\bar{\psi}\psi]}_{\text{free fermion}} - \underbrace{[g\bar{\psi}\gamma^\mu\psi A_\mu]}_{\text{interaction}} + \underbrace{[-\frac{1}{16\pi} F^{\mu\nu} F_{\mu\nu}]}_{\text{free gauge particle}}$$

The existence of the photon, coupling to fermions in the usual way, guarantees local gauge invariance.

Another (conventional) way of describing this...

$$\psi \rightarrow e^{-ig\lambda(x)} \psi$$

implies

$$\partial_\mu \psi \rightarrow e^{-ig\lambda(x)} [\partial_\mu - ig\partial_\mu \lambda] \psi$$

this term destroys the invariance of the derivative term in  $\mathcal{L}$

solution:

In  $\mathcal{L}$ , replace  $\partial_\mu$  by

$$D_\mu \equiv \partial_\mu + igA_\mu$$

"covariant derivative"

then

$$D_\mu \psi \rightarrow e^{-ig\lambda(x)} D_\mu \psi$$

This is equivalent to adding  $-g\bar{\psi}\gamma^\mu\psi A_\mu$  to  $\mathcal{L}$ .

# Yang Mills Theory (Weak Interactions)

W couples 2 different types of fermions together.

⇒ Start with a two fermion theory

(for simplicity, use the same mass for each)

$$\mathcal{L} = i \bar{\psi}_1 \gamma^\mu \partial_\mu \psi_1 - m \bar{\psi}_1 \psi_1 + i \bar{\psi}_2 \gamma^\mu \partial_\mu \psi_2 - m \bar{\psi}_2 \psi_2$$

rewrite as  $\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}$   $\bar{\psi} = (\bar{\psi}_1 \bar{\psi}_2)$

$$\mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi \leftarrow 2 \text{ vector}$$

Previously we used  $\psi \rightarrow e^{-i\theta^2} \psi$

↑ phase transformation  
(a.k.a.  $1 \times 1$  unitary)

For a 2 fermion theory use:

$$\psi \rightarrow U \psi \quad U U^\dagger = 1$$

where  $U$  is a  $2 \times 2$  unitary matrix ( $\in SU(2)$ )

This corresponds to rotations btwn  $\psi_1$  &  $\psi_2$   
(rotations in "weak isospin" space)

A unitary matrix can be written as the exponential of some Hermitian matrix.

$$U = e^{iH} \quad \text{with} \quad H^\dagger = H$$

and a  $2 \times 2$  Hermitian matrix can be written

$$H = \theta I + a \cdot \sigma \leftarrow \text{Pauli matrices}$$

Thus,

$$U = e^{i\theta} e^{i\sigma \cdot a}$$

We already looked at this. It leads to the photon.

$\uparrow$   $\uparrow$  Now look at this.

$$\psi \rightarrow e^{-ig \sigma \cdot \lambda(x)} \psi \equiv S \psi$$

$\uparrow$   $\lambda$  is now a 3 vector

The derivative  $\partial_\mu \psi \rightarrow S \partial_\mu \psi + (\partial_\mu S) \psi$  is not invariant. Fix it by replacing with the covariant derivative.

$$D_\mu \equiv \partial_\mu + ig \sigma \cdot A_\mu$$

where the  $A_\mu$  transform  $\uparrow$  3 new fields in such a way that  $\mathcal{L}$  is invariant.

In the case of small  $\lambda$  (see Griffiths 11.58-11.61)

$$A_\mu \rightarrow A_\mu + \partial_\mu \lambda + 2g(\lambda \times A_\mu)$$

The case of finite  $\lambda$  is no more illuminating so we will stick to this for simplicity.

$$\mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi$$

$$= i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi - g \bar{\psi} \gamma^\mu \sigma \psi \cdot A_\mu$$

is now invariant.

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Since we introduced 3 new gauge fields we should include their own free Lagrangian.

$$\mathcal{L}_A = \frac{-1}{16\pi} F_1^{\mu\nu} F_{\mu\nu 1} - \frac{1}{16\pi} F_2^{\mu\nu} F_{\mu\nu 2} - \frac{1}{16\pi} F_3^{\mu\nu} F_{\mu\nu 3}$$

$$= \frac{-1}{16\pi} F^{\mu\nu} \cdot F_{\mu\nu} \leftarrow F_{\mu\nu} \text{ is now a 3 vector}$$

### PROBLEM

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu \text{ is not invariant!}$$

$$A_\mu \rightarrow A_\mu + \partial_\mu \lambda + 2g (\lambda \times A_\mu)$$

leads to

$$F^{\mu\nu} \rightarrow F^{\mu\nu} + 2g [\lambda \times F^{\mu\nu} + A^\mu \times \partial^\nu \lambda - A^\nu \times \partial^\mu \lambda]$$

and

$$F^{\mu\nu} \cdot F_{\mu\nu} \rightarrow F^{\mu\nu} \cdot F_{\mu\nu} + 8g (A_\mu \times F^{\mu\nu}) \cdot \partial_\mu \lambda$$

To fix this

$$F^{\mu\nu} \equiv \partial^\mu A^\nu - \partial^\nu A^\mu - 2g (A^\mu \times A^\nu)$$

then

$$F^{\mu\nu} \rightarrow F^{\mu\nu} + 2g (\lambda \times F^{\mu\nu})$$

and

$$F^{\mu\nu} \cdot F_{\mu\nu} \rightarrow F^{\mu\nu} \cdot F_{\mu\nu} \text{ to 1st order in } \lambda$$

Conclusion

$$\mathcal{L} = \underbrace{i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi}_{2 \text{ fermions}} - \underbrace{\frac{1}{16\pi} F^{\mu\nu} \cdot F_{\mu\nu}}_{3 \text{ massless vector fields}} - \underbrace{g \bar{\psi} \gamma^\mu \psi \cdot A_\mu}_{\text{fermion gauge interactions}}$$

# Chromodynamics (short form)

start with:

$$\mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi \quad \text{where} \quad \psi = \begin{pmatrix} \psi_r \\ \psi_b \\ \psi_g \end{pmatrix}$$

$$\psi \rightarrow U \psi \quad \text{with } U \text{ } 3 \times 3 \text{ unitary } (\in SU(3))$$

$$U = e^{iH} \quad \text{with } H = \Theta I + a \cdot \lambda \quad \leftarrow \text{8 Gell-Mann matrices}$$

$$\text{we consider } \psi \rightarrow e^{-ig \lambda \cdot \phi(x)} \psi$$

$$D_\mu \equiv \partial_\mu + ig \lambda \cdot A_\mu \quad \leftarrow \text{8 new fields}$$

$$A_\mu \rightarrow A_\mu + \partial_\mu \phi + 2g (\phi \times A_\mu)$$

$$\text{where } (\phi \times A_\mu)_k = \sum_{i,j} f_{ijk} \phi_i A_{\mu j} \quad \leftarrow \text{structure consts.}$$

$$F^{\mu\nu} \equiv \partial^\mu A^\nu - \partial^\nu A^\mu - 2g (A^\mu \times A^\nu)$$

$$\mathcal{L} = i \bar{\psi} \gamma^\mu \partial_\mu \psi - m \bar{\psi} \psi - \frac{1}{16\pi} F^{\mu\nu} \cdot F_{\mu\nu} - g \bar{\psi} \gamma^\mu \lambda \psi \cdot A_\mu$$

# From $\mathcal{L}$ to Feynman Rules

## Propagators

Euler-Lagrange yields free field equations

KG  $[\rho^2 - m^2] \phi = 0$  where we used

Dirac  $[\not{\rho} - m] \psi = 0$   $i \partial_\mu = p_\mu$

Proca  $[(-\rho^2 + m^2) g_{\mu\nu} + p_\mu p_\nu] A^\nu = 0$

The propagator is  $i$  times the inverse of the term in  $[\ ]$  brackets.

$$\frac{i}{\rho^2 - m^2} \quad \text{spin } 0$$

$$\frac{i}{\not{\rho} - m} = \frac{i(\not{\rho} + m)}{\rho^2 - m^2} \quad \text{spin } \frac{1}{2}$$

$$\frac{-i}{\rho^2 - m^2} \left[ g_{\mu\nu} - \frac{p_\mu p_\nu}{m^2} \right] \quad \text{spin } 1$$

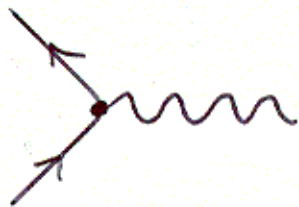
To understand why this is the case requires a knowledge of Green's functions.

## Vertex factors

$$i \mathcal{L}_{\text{INT}}^{\text{QED}} = -ig \bar{\psi} \gamma^\mu \psi A_\mu$$

describes

interaction  
of 3 fields



with vertex factor

$$-ig \gamma^\mu$$

$$g \equiv -g_{e.m.}$$

Similarly,

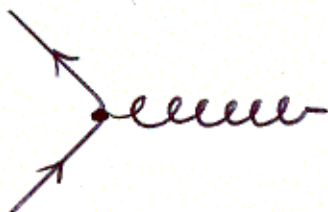
$$i \mathcal{L}_{\text{INT}}^{\text{QCD}} = -ig \bar{\psi} \gamma^\mu \lambda \psi \cdot A_\mu$$

describes

with vertex factor

$$-ig \gamma^\mu \lambda$$

$$g \equiv \frac{g_s}{2}$$



( $g_s$  traditionally defined with an extra factor of 2)

For QCD, there are also interaction terms in  $F^{\mu\nu} \cdot F_{\mu\nu}$

$$F^{\mu\nu} \cdot F_{\mu\nu} = (\partial^\mu A^\nu - \partial^\nu A^\mu) \cdot (\partial_\mu A_\nu - \partial_\nu A_\mu) \\ - 2g [(\partial^\mu A^\nu - \partial^\nu A^\mu) \cdot (A_\mu \times A_\nu) + (A^\mu \times A^\nu) \cdot (\partial_\mu A_\nu - \partial_\nu A_\mu)] \\ + 4g^2 [(A^\mu \times A^\nu) \cdot (A_\mu \times A_\nu)]$$

Line 2 describes interaction of 3 fields



Line 3 describes interaction of 4 fields





Consider two <sup>REAL</sup> scalar fields ...

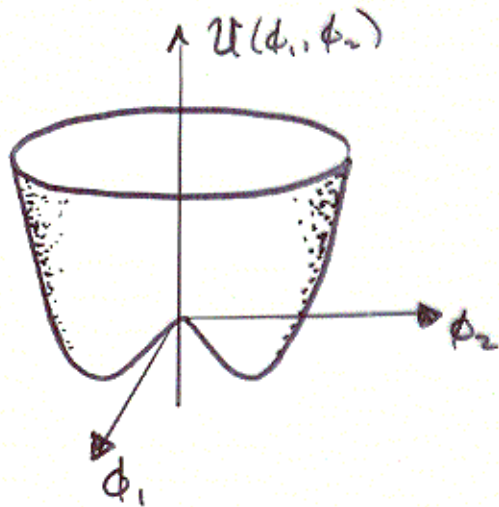
$$\mathcal{L} = \frac{1}{2}(\partial_\mu \phi_1)(\partial^\mu \phi_1) + \frac{1}{2}(\partial_\mu \phi_2)(\partial^\mu \phi_2) + \frac{1}{2}\mu^2(\phi_1^2 + \phi_2^2) - \frac{\lambda}{4}(\phi_1^2 + \phi_2^2)^2$$

choose a minimum at

$$\phi_1 = \mu/\lambda \quad \phi_2 = 0$$

and use change of variable

$$\eta = \phi_1 - \mu/\lambda \quad \xi = \phi_2$$



massive scalar

massless scalar (Goldstone boson)

$$\mathcal{L} = \left[ \frac{1}{2}(\partial_\mu \eta)(\partial^\mu \eta) - \mu^2 \eta^2 \right] + \left[ \frac{1}{2}(\partial_\mu \xi)(\partial^\mu \xi) \right] + \mu\lambda(\eta^3 + \eta\xi^2) - \frac{\lambda^2}{4}(\eta^4 + \xi^4 + 2\eta^2\xi^2) + \mu^2/4\lambda^2$$

This is called spontaneously broken symmetry. The ground state  $\phi$  does not preserve the symmetry of  $\mathcal{L}$ .

# The Higgs Mechanism

First, a change of notation

$$\phi = \phi_1 + i\phi_2 \quad \phi^* = \phi_1 - i\phi_2$$

$$\mathcal{L} = \frac{1}{2}(\partial_\mu \phi)^*(\partial^\mu \phi) + \frac{1}{2}\mu^2 \phi^* \phi - \frac{1}{4}\lambda^2(\phi^* \phi)^2$$

$\mathcal{L}$  is invariant under a global phase shift

$$\phi \rightarrow e^{i\theta} \phi$$

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To make local invariance we require

$$\partial_\mu \rightarrow \mathcal{D}_\mu = \partial_\mu + ig A_\mu \quad (\text{covariant derivative})$$

The resulting Lagrangian:

$$\mathcal{L} = \frac{1}{2}(\partial_\mu - ig A_\mu)\phi^*(\partial_\mu + ig A_\mu)\phi + \frac{1}{2}\mu^2 \phi^* \phi - \frac{1}{4}\lambda^2(\phi^* \phi)^2 - \frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu}$$

is invariant.

Since  $\mathcal{L}$  is invariant under  $\phi \rightarrow e^{i\Theta(x)}\phi$   
 we can transform

$$\phi = \phi_1 + i\phi_2 \rightarrow (\cos\Theta + i\sin\Theta)(\phi_1 + i\phi_2)$$

and choose  $\Theta (-\tan^{-1} \phi_2/\phi_1)$  such that  
 $\phi' = \phi'_1$  is purely real and  $\phi'_2 = 0$ .

Once again, change variables

$$\eta = \phi'_1 - \mu/\lambda \quad \xi = \phi'_2 = 0$$

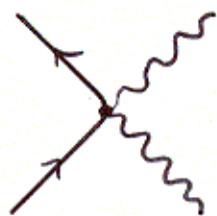
Result:

$$\begin{aligned} \mathcal{L} = & \frac{1}{2} \partial_\mu \eta \partial^\mu \eta - \mu^2 \eta^2 && \leftarrow \text{massive scalar} \\ & + \frac{1}{16\pi} F^{\mu\nu} F_{\mu\nu} + \frac{1}{2} g^2 \frac{\mu^2}{\lambda^2} A_\mu A^\mu && \leftarrow \text{massive vector!} \\ & + \frac{\mu}{\lambda} g^2 \eta A_\mu A^\mu + \frac{1}{2} g^2 \eta^2 A_\mu A^\mu - \lambda \mu \eta^3 - \frac{1}{4} \lambda^2 \eta^4 \\ & + \frac{1}{4} \mu^4 / \lambda^2 && \leftarrow \text{constant} \end{aligned}$$

↑ interactions

Where did the mass for  $A_\mu$  come from?

In the original  $\mathcal{L}$  we had the term:



$$g^2 A_\mu A^\mu \phi^* \phi$$

Under SSB  $\phi \rightarrow \eta + \mu/\lambda$ ,  
giving rise to the term  $g^2 A_\mu A^\mu \frac{\mu^2}{\lambda^2}$ ,  
which describes a vector mass.

SSB of a scalar that couples to a vector field can produce a mass term for the vector field.

In the SM we start with a weak isospin doublet of Higgs particles.

Under symmetry breaking, this gives one massive scalar field  $\dagger$  3 Goldstones

If these are coupled to the four electroweak vector bosons in a locally gauge invariant way under  $SU(2) \dagger U(1)$  transformations...

3 GBs  $\Rightarrow$  3 massive vector bosons

By the way, it is also possible to start with massless fermions that couple to the Higgs via terms like  $-g \bar{\psi} \psi \phi$ . When SSB occurs these terms give mass terms  $-g \bar{\psi} \psi \frac{\mu}{\lambda}$ .

For full details see "Quarks  $\dagger$  Leptons" Halzen  $\dagger$  Martin